

Superintegrable classical Zernike system

George S. Pogosyan, Kurt Bernardo Wolf, and Alexander Yakhno

Citation: *Journal of Mathematical Physics* **58**, 072901 (2017); doi: 10.1063/1.4990793

View online: <http://dx.doi.org/10.1063/1.4990793>

View Table of Contents: <http://aip.scitation.org/toc/jmp/58/7>

Published by the [American Institute of Physics](#)

Articles you may be interested in

[Quantum superintegrable Zernike system](#)

Journal of Mathematical Physics **58**, 072101 (2017); 10.1063/1.4990794

[Fock representations of Q-deformed commutation relations](#)

Journal of Mathematical Physics **58**, 073501 (2017); 10.1063/1.4991671

[Hartree-Fock theory with a self-generated magnetic field](#)

Journal of Mathematical Physics **58**, 062108 (2017); 10.1063/1.4987022

[Exactly solvable Hermite, Laguerre, and Jacobi type quantum parametric oscillators](#)

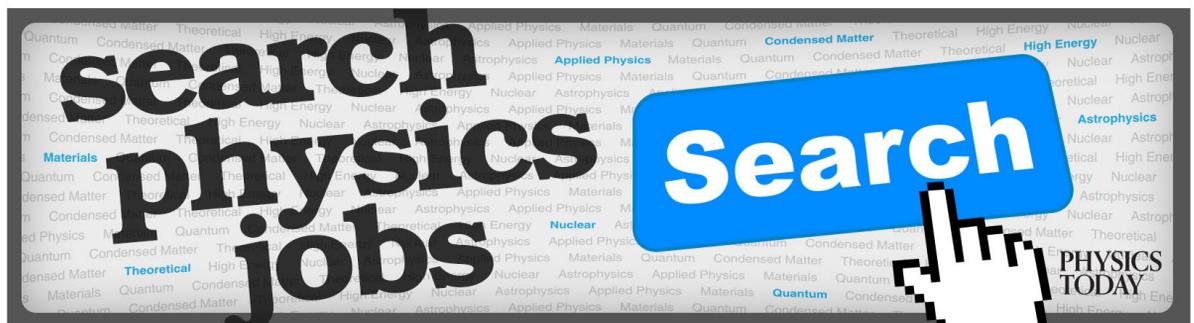
Journal of Mathematical Physics **57**, 122107 (2016); 10.1063/1.4972293

[Transition operators](#)

Journal of Mathematical Physics **58**, 051703 (2017); 10.1063/1.4983479

[Pedal coordinates, dark Kepler, and other force problems](#)

Journal of Mathematical Physics **58**, 063505 (2017); 10.1063/1.4984905



Superintegrable classical Zernike system

George S. Pogosyan,^{1,2} Kurt Bernardo Wolf,³ and Alexander Yakhno¹

¹*Departamento de Matemáticas, Centro Universitario de Ciencias Exactas e Ingenierías, Universidad de Guadalajara, Guadalajara, Mexico*

²*Yerevan State University, Yerevan, Armenia and Joint Institute for Nuclear Research, Dubna, Russia*

³*Instituto de Ciencias Físicas, Universidad Nacional Autónoma de México, Cuernavaca, Mexico*

(Received 13 February 2017; accepted 16 June 2017; published online 6 July 2017)

We consider the differential equation that Zernike proposed to classify aberrations of wavefronts in a circular pupil, as if it were a classical Hamiltonian with a non-standard potential. The trajectories turn out to be closed ellipses. We show that this is due to the existence of higher-order invariants that close into a cubic Higgs algebra. The Zernike classical system thus belongs to the class of superintegrable systems. Its Hamilton-Jacobi action separates into three vertical projections of polar coordinates of sphere, polar, and equidistant coordinates on half-hyperboloids, and also in elliptic coordinates on the sphere. *Published by AIP Publishing.* [<http://dx.doi.org/10.1063/1.4990793>]

I. INTRODUCTION: THE ZERNIKE OPERATOR

In Ref. 27 (p. 700), Frits Zernike proposed a two-dimensional differential equation whose polynomial solutions provide an orthogonal basis for functions $f(\mathbf{r})$ in a Hilbert space $\mathcal{L}_Z^2(\mathcal{D}_1)$ over the unit disk $\mathbf{r} \in \mathcal{D}_1$, $|\mathbf{r}| \leq 1$ which—importantly—have a constant absolute value on the boundary circle: $|f(\mathbf{r})|_{|\mathbf{r}|=1} = 1$. This Zernike basis is thus distinct from the well-known bases of Bessel functions over the disk whose values (or logarithmic derivatives) vanish on a boundary circle. The differential operator and eigenvalue equation of Zernike are

$$\hat{Z}^{(\alpha,\beta)}f(\mathbf{r}) := \left(\nabla^2 + \alpha(\mathbf{r} \cdot \nabla)^2 + \beta \mathbf{r} \cdot \nabla \right) f(\mathbf{r}) = -E f(\mathbf{r}). \quad (1)$$

The requirement that this operator should be self-adjoint under the inner product $(f_1, f_2)_{\mathcal{D}_1} := \int_{\mathcal{D}_1} d^2\mathbf{r} f_1(\mathbf{r})^* f_2(\mathbf{r})$, i.e., $(\hat{Z}f_1, f_2)_{\mathcal{D}_1} = (f_1, \hat{Z}f_2)_{\mathcal{D}_1}$, constrains the coefficients to have the values $(\alpha_Z, \beta_Z) := (-1, -2)$.²⁷ In this paper however, we let α and β take arbitrary real values, to be later constrained to those regions that lead to the closed orbits that we consider to be the main feature of interest of the Zernike system.

For $\hat{Z}^{(\alpha_Z, \beta_Z)}$ in (1), the polar factored solutions $Z_{n,m}(r) \exp(im\phi)$, $|m| \leq n$, correspond to the eigenvalues $E = n(n+2)$; when normalized to $Z_{n,m}(1) = 1$, the radial functions are the Zernike polynomials.²⁷ These can be related to the Jacobi polynomials $\sim P_n^{(m-n, 0)}(2r^2 - 1)$ whose interval of orthogonality is $|-1 \leftrightarrow r|_0^1$. It was remarked in Ref. 4 that the reasons for postulating Eq. (1) were rather arbitrary, so its authors used the Gram-Schmidt method to find the same polynomial solutions from first principles. Zernike polynomials have wide applications in the correction of optical aberrations by describing wavefronts at circular pupils (see, for example, Ref. 5); they also display a host of enticing mathematical properties^{10,11,15,20,22,26} that are characteristic of algebraic structures.

When $\alpha = 0$, $\hat{Z}^{(0,\beta)}$ reduces to a linear combination of generators of the real symplectic algebra $\mathfrak{sp}(4, \mathbb{R})$ under Poisson brackets or commutators (Ref. 25, Sec. 11.4); also when $\beta = 0$, then (1) becomes simply the Laplace equation with plane wave solutions $\sim \exp(i\mathbf{k} \cdot \mathbf{r})$, $|\mathbf{k}|^2 = E$ or, adapted to polar coordinates (r, ϕ) , multipole solutions $\sim J_m(kr) e^{im\phi}$ with Bessel functions, where the radial wavenumber k may or may not be quantized according to whether the boundary conditions are set at a finite or infinite radius. On the other hand, when $\alpha \neq 0$ but $\beta = 0$, the Zernike equation (1) reduces to the kinetic part of a nonlinear oscillator Hamiltonian.⁶ We shall keep their generic values $(\alpha, \beta) \in \mathcal{R}^2$ and particularize when convenient.

We found that it is of interest to examine the *classical counterpart* of the Zernike system, which in “wave” (or quantum mechanical) form is (1). The process of *de-quantization* of this equation consists in replacing

$$\nabla \mapsto i\mathbf{p} = i \begin{pmatrix} p_x \\ p_y \end{pmatrix}, \quad \mathbf{r} = \begin{pmatrix} x \\ y \end{pmatrix}, \quad r := |\mathbf{r}|, \quad (2)$$

$$\nabla^2 \mapsto -(p_x^2 + p_y^2) = -\left(p_r^2 + \frac{p_\phi^2}{r^2}\right), \quad \mathbf{r} \cdot \nabla \mapsto i(xp_x + yp_y) = i r p_r. \quad (3)$$

The operator (1) thus yields a classical Hamiltonian function $H^{(\alpha,\beta)} = -\hat{Z}^{(\alpha,\beta)}$ which depends on two coordinates and two momenta. In Cartesian and polar coordinates, it is

$$H^{(\alpha,\beta)} := (p_x^2 + p_y^2) + \alpha(xp_x + yp_y)^2 - i\beta(xp_x + yp_y) \quad (4)$$

$$= (1 + \alpha r^2)p_r^2 + p_\phi^2/r^2 - i\beta r p_r, \quad (5)$$

and its value is the energy E . The appearance of $i = \sqrt{-1}$ in this Hamiltonian seems indeed anomalous, yet our calculations will show that at the end we have a purely real classical system whose trajectories can be found explicitly. We have been advised on a possible relation of the Zernike system with other non-hermitian Hamiltonian systems with complex potentials in the context of PT-invariant theories (see, for instance, Refs. 1 and 2), although in its quantum version (following paper), the Zernike Hamiltonian *is* self-adjoint, albeit on a domain that is not usual in quantum mechanics.

The Hamilton-Jacobi method is particularly apt to solve this system, where we shall preferentially use the polar coordinates (r, ϕ) and their momenta (p_r, p_ϕ) in (5). Since $H^{(\alpha,\beta)} = E$ is independent of time and the angular coordinate ϕ is cyclic, the *action* function $S(r, \phi)$ (also called Hamilton’s *principal* function) that satisfies the Hamilton-Jacobi equation $H + \partial S/\partial t = 0$ can be separated in the form

$$S(r, \phi) = R(r) + p_\phi \phi - Et. \quad (6)$$

The space derivatives of this function yield the polar momenta p_r and p_ϕ as

$$p_r = \frac{\partial S(r, \phi)}{\partial r}, \quad r = -\frac{\partial S(r, \phi)}{\partial p_r}, \quad p_\phi = \frac{\partial S(r, \phi)}{\partial \phi}, \quad \phi = -\frac{\partial S(r, \phi)}{\partial p_\phi}. \quad (7)$$

In Sec. II, we shall use the derivatives of (6) with respect to the radius r and the angle ϕ , to find the *geometric* trajectories $r(\phi)$, which are closed ellipses. Then in Sec. III, the *dynamical* trajectories $\mathbf{r}(t)$ will be found differentiating the action $S(r, \phi)$ with respect to the energy. The symmetries behind the closure of the orbits will be elucidated in Sec. IV, where Eq. (1) is separated into three spherical, six hyperbolic, and elliptic coordinates, and shown to lead to constants of motion. In Sec. V, we show that the operators which characterize these constants close into a cubic superintegrable algebra and offer some additional comments.

II. GEOMETRIC TRAJECTORIES $r(\phi)$

The derivative of the action function (6) with respect to the radius r is the radial momentum,

$$p_r = \frac{\partial S(r, \phi)}{\partial r} = \frac{\partial R(r)}{\partial r}. \quad (8)$$

Replacing p_r in (5) yields a quadratic algebraic equation for the derivative of $R(r)$, namely,

$$(1 + \alpha r^2) \left(\frac{\partial R(r)}{\partial r} \right)^2 - i\beta r \left(\frac{\partial R(r)}{\partial r} \right) + \frac{p_\phi^2}{r^2} = E, \quad (9)$$

whose two solutions are

$$\frac{\partial R(r)}{\partial r} = \frac{i\beta r \pm \sqrt{-\beta^2 r^2 - 4(1 + \alpha r^2)(p_\phi^2/r^2 - E)}}{2(1 + \alpha r^2)}. \quad (10)$$

From here we find $R(r)$ through the indefinite integral

$$R(r) = \int dr \left(\frac{i\beta r}{2(1 + \alpha r^2)} \pm \frac{\sqrt{(\alpha E - \frac{1}{4}\beta^2)r^2 + (E - \alpha p_\phi^2) - p_\phi^2/r^2}}{1 + \alpha r^2} \right). \tag{11}$$

We can now find the trajectories that relate r and ϕ by differentiating (6) with respect to p_ϕ ,

$$\frac{\partial S(r, \phi)}{\partial p_\phi} = \frac{\partial R(r)}{\partial p_\phi} + \phi = \phi_o, \tag{12}$$

where ϕ_o is a constant of the motion given by the initial conditions. The derivative of $R(r)$ in (11) with respect to p_ϕ is then

$$\frac{\partial R(r)}{\partial p_\phi} = \pm \int dr \frac{\partial}{\partial p_\phi} \frac{\sqrt{(\alpha E - \frac{1}{4}\beta^2)r^2 - (\alpha + 1/r^2)p_\phi^2 + E}}{1 + \alpha r^2} \tag{13}$$

$$= \mp p_\phi \int \frac{dr}{r} \frac{1}{\sqrt{(\alpha E - \frac{1}{4}\beta^2)r^4 + (E - \alpha p_\phi^2)r^2 - p_\phi^2}} \tag{14}$$

$$= \mp \frac{p_\phi}{2} \int dz \frac{1}{z\sqrt{a + bz + cz^2}}, \tag{15}$$

where in the last equality, we have substituted $z = r^2$ with $dr/r = \frac{1}{2}dz/z$, and we define

$$a := -p_\phi^2, \quad b := E - \alpha p_\phi^2, \quad c := \alpha E - \frac{1}{4}\beta^2. \tag{16}$$

We note that the imaginary summand in (11) is absent from this equation and thus from the system. The double sign in (13) corresponds to the $\pm p_\phi$ angular momentum of a trajectory traversed in opposite directions.

One finds the indefinite integral solved in Ref. 8, Eq. (2.266), with various expressions involving inverse trigonometric and hyperbolic functions, or logarithms, depending on the signs of the constants; in our case (16) $a < 0$ and for $b^2 - 4ac = (E + \alpha p_\phi^2)^2 - \beta^2 p_\phi^2 > 0$, the integral is

$$\int dz \frac{1}{z\sqrt{a + bz + cz^2}} = \frac{1}{\sqrt{-a}} \arcsin \frac{2a + bz}{z\sqrt{b^2 - 4ac}}. \tag{17}$$

Thus, joining Eqs. (12) (16), and (17), we obtain

$$\phi - \phi_o = -\frac{\partial R(r)}{\partial p_\phi} = \frac{1}{2} \arcsin \frac{(E - \alpha p_\phi^2)r^2 - 2p_\phi^2}{r^2 \sqrt{(E + \alpha p_\phi^2)^2 - \beta^2 p_\phi^2}}, \tag{18}$$

and this leads to $\phi(r^2)$ in the form

$$\sin 2(\phi - \phi_o) = \frac{Ar^2 - B}{Cr^2}, \quad \begin{cases} A := E - \alpha p_\phi^2, \\ B := 2p_\phi^2, \\ C := \sqrt{(E + \alpha p_\phi^2)^2 - \beta^2 p_\phi^2}. \end{cases} \tag{19}$$

We can invert the dependence to $r(\phi)$ by solving for the square radius and setting for convenience $\phi_o = -\frac{1}{4}\pi$,

$$r^2(\phi) = \frac{B}{A - C \cos 2\phi} = \frac{2p_\phi^2}{(E - \alpha p_\phi^2) - \sqrt{(E + \alpha p_\phi^2)^2 - \beta^2 p_\phi^2} \cos 2\phi} \tag{20}$$

$$= \frac{D}{1 - \varepsilon \cos 2\phi}, \quad \begin{cases} D := B/A = 2p_\phi^2/(E - \alpha p_\phi^2), \quad E \neq \alpha p_\phi^2, \\ \varepsilon := \frac{C}{A} = \frac{\sqrt{(E + \alpha p_\phi^2)^2 - \beta^2 p_\phi^2}}{E - \alpha p_\phi^2}. \end{cases} \tag{21}$$

This is the parametric equation for ellipses, provided that

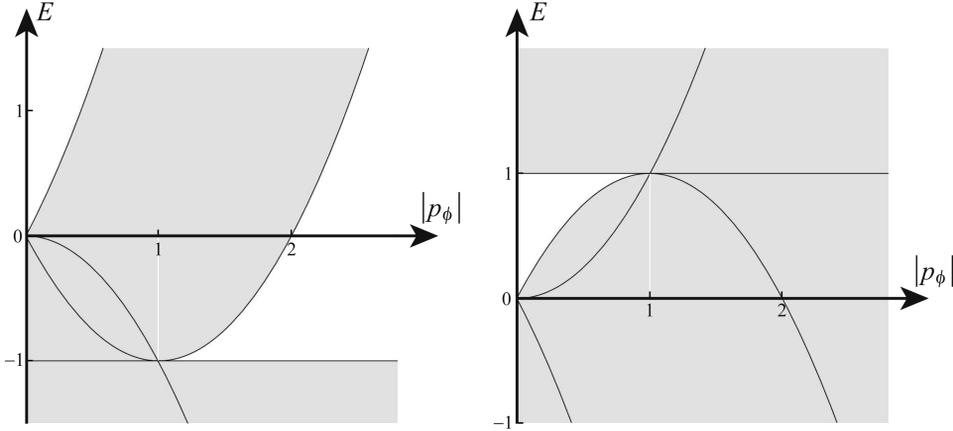


FIG. 1. Regions of the plane of angular momentum p_ϕ and energy E , where closed trajectories are allowed by the inequalities (22) (in white) for (α, β) -Zernike systems. *Left*: Allowed regions for the original Zernike system $(\alpha_Z, \beta_Z) = (-1, -2)$. *Right*: Allowed regions for $(\alpha, \beta) = (1, 2)$. Closed elliptical trajectories do not occur in the gray regions. The structure of these regions is generic for all α and $\beta \neq 0$. The units of E in these graphs are $\beta^2/4|\alpha|$ and the units of $|p_\phi|$ are $\frac{1}{2}|\beta/\alpha|$.

$$\begin{aligned}
 \varepsilon \text{ real} &\Rightarrow C^2 \geq 0 \Rightarrow \begin{cases} E \leq -\alpha p_\phi^2 - |\beta p_\phi|, \\ E \geq -\alpha p_\phi^2 + |\beta p_\phi|, \end{cases} \\
 |\varepsilon| < 1 &\Rightarrow A^2 > C^2 \Rightarrow 4\alpha E < \beta^2, \\
 r^2(\phi) > 0 &\Rightarrow D > 0 \Rightarrow E > \alpha p_\phi^2.
 \end{aligned} \tag{22}$$

These conditions restrict the range of energies E and angular momenta p_ϕ where the trajectories are real and closed. As shown in Fig. 1 (left) for the generic Zernike range $\alpha < 0$, $\beta \neq 0$, the first condition *excludes* the energy interval between the two parabolas, $-\alpha p_\phi^2 - |\beta p_\phi| \leq E \leq -\alpha p_\phi^2 + |\beta p_\phi|$; the second inequality is (for $\alpha < 0$) a *lower bound* $E > -\beta^2/4|\alpha|$ (equal to -1 for the Zernike case); lastly, the third condition excludes the interior of the parabola $E = \alpha p_\phi^2$ that has its apex at the origin, and which eliminates the region $|p_\phi| < -\frac{1}{2}|\beta|/|\alpha|$ that was (left) allowed by the previous two conditions.

In Fig. 1 (right), we show the allowed regions for the generic Zernike range $\alpha > 0$, $\beta \neq 0$. The two parabolas stemming from the first inequality in (22) under $\alpha \leftrightarrow -\alpha$ reflect the E -axis; the second inequality in (22) is now the *upper bound* $E < \beta^2/4\alpha$; and the third inequality allows elliptic orbits in the remaining interior of the parabola, namely, $-\alpha p_\phi^2 + |\beta p_\phi| < E < \beta^2/4\alpha$ for $0 \leq |p_\phi| < |\beta|/2\alpha$. Finally, when $\alpha = 0$, the “forbidden” region between the two parabolas due to the first condition in (22) becomes $-\beta p_\phi \leq E \leq \beta p_\phi$, while the second two conditions are satisfied by $E > 0$, so that closed elliptical trajectories occur for all $E \geq |\beta p_\phi|$.

Since we took $\phi_o = -\frac{1}{4}\pi$, the y -axis is at $\phi = 0$ and the x -axis at $\phi = \frac{1}{2}\pi$. The semi-major and semi-minor axes of the ellipse are, respectively,

$$\mu_y := \sqrt{\frac{D}{1-\varepsilon}} = \sqrt{\frac{B}{A-C}}, \quad \mu_x := \sqrt{\frac{D}{1+\varepsilon}} = \sqrt{\frac{B}{A+C}}. \tag{23}$$

The *area* of this ellipse is given by π times the product of the two semi-axes,

$$\text{area} = \pi \mu_x \mu_y = \frac{\pi D}{\sqrt{1-\varepsilon^2}} = \frac{\pi B}{\sqrt{A^2 - C^2}} = \frac{2\pi |p_\phi|}{\sqrt{\beta^2 - 4\alpha E}}. \tag{24}$$

III. DYNAMICAL TRAJECTORIES $r(t)$ AND ORBITS

We return now to the integral expression for $R(r)$ in (11), differentiating the action $S(r, \phi)$ in (6) now with respect to the energy E ,

$$\frac{\partial S(r, \phi)}{\partial E} = \frac{\partial R(r)}{\partial E} - t = -t_o, \quad (25)$$

where t_o is the initial time constant. Instead of (13)–(15), we now have

$$\frac{\partial R(r)}{\partial E} = \pm \int dr \frac{\partial}{\partial E} \frac{\sqrt{(\alpha E - \frac{1}{4}\beta^2)r^2 - (\alpha + 1/r^2)p_\phi^2 + E}}{1 + \alpha r^2} \quad (26)$$

$$= \pm \frac{1}{2} \int dr \frac{1}{\sqrt{(\alpha E - \frac{1}{4}\beta^2)r^2 - (\alpha + 1/r^2)p_\phi^2 + E}} \quad (27)$$

$$= \pm \frac{1}{4} \int dz \frac{1}{\sqrt{a + bz + cz^2}}, \quad (28)$$

where as before we have set $z = r^2$, and a, b, c are again given by (16). The indefinite integral can be found in Ref. 8, Eq. (2.261), it is

$$\int dz \frac{1}{\sqrt{a + bz + cz^2}} = \frac{-1}{\sqrt{-c}} \arcsin \frac{2cz + b}{\sqrt{b^2 - 4ac}}. \quad (29)$$

The conditions for this integral to be proper, $c < 0$ and $b^2 - 4ac > 0$, also lead to (22), while the solutions corresponding to (19) are now

$$\sin(4(t - t_o)\sqrt{U}) = \frac{A - 2U r(t)^2}{C}, \quad U := \frac{1}{4}\beta^2 - \alpha E = \frac{A^2 - C^2}{2B} > 0, \quad (30)$$

with A and C given by (19).

From here we can extract the dependence of the square radius of the trajectory on time as (21) did for the angle. We choose t_o such that $r(t)|_{t=0} = \mu_y$ is the semi-major axis in (23), i.e., $4t_o\sqrt{U} = \frac{1}{2}\pi$, so $t_o = \frac{1}{8}\pi/\sqrt{U}$, and write

$$\begin{aligned} r^2(t) &= \frac{A + C \cos(4t\sqrt{U})}{2U} \\ &= \frac{E + \sqrt{(E + \alpha p_\phi^2)^2 - \beta^2 p_\phi^2} \cos(2t\sqrt{\beta^2 - 4\alpha E}) - \alpha p_\phi^2}{\frac{1}{2}\beta^2 - 2\alpha E}. \end{aligned} \quad (31)$$

This is a periodic function of time, with period $4T\sqrt{U} = 2\pi$ or

$$T = \pi \sqrt{\beta^2 - 4\alpha E}. \quad (32)$$

In the generalized Zernike range $\alpha < 0$, the radicand is positive; when $\alpha > 0$, the second inequality in (22) prevents the orbits from being closed for $\alpha E > \frac{1}{4}\beta^2$. Although orbits in the Zernike range are ellipses, they *differ* from the isochronous orbits of the classical harmonic oscillator, whose period does not depend on their energy.⁷

As a function of time, the trajectories $(x(t), y(t))$ can be found from the previous expressions, (21) and (31), as

$$\begin{aligned} x(t) &= r \sin \phi = r \sqrt{\frac{1}{2}(1 - \cos 2\phi)} = \frac{1}{\sqrt{2\varepsilon}} \sqrt{(\varepsilon - 1)r^2(t) + D} \\ &= \frac{1}{\sqrt{2\varepsilon}} \sqrt{(\varepsilon - 1) \frac{A + C \cos(4t\sqrt{U})}{2U} + D}, \end{aligned} \quad (33)$$

$$\begin{aligned} y(t) &= r \cos \phi = r \sqrt{\frac{1}{2}(1 + \cos 2\phi)} = \frac{1}{\sqrt{2\varepsilon}} \sqrt{(\varepsilon + 1)r^2(t) - D} \\ &= \frac{1}{\sqrt{2\varepsilon}} \sqrt{(\varepsilon + 1) \frac{A + C \cos(4t\sqrt{U})}{2U} - D}, \end{aligned} \quad (34)$$

and are shown in Fig. 2 for the Zernike case $(\alpha_Z, \beta_Z) = (-1, -2)$ but are valid for the range $\alpha < 0$.

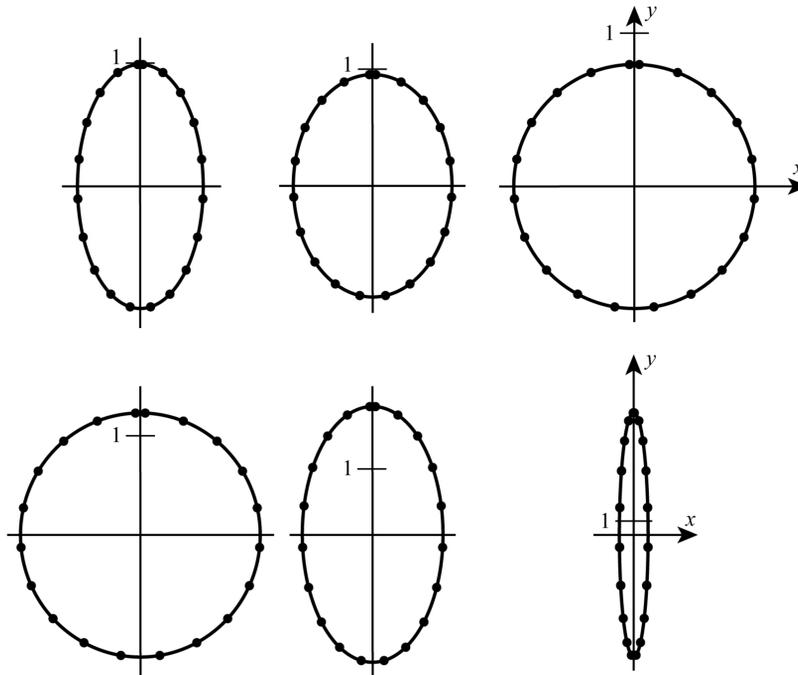


FIG. 2. Trajectories $(x(t), y(t))$ in the classical Zernike system $(\alpha_Z, \beta_Z) = (-1, -2)$ and angular momentum $p_\phi = 3$, for equidistant times $t \in [0, T]$. *Upper row*: trajectories inside the disk \mathcal{D}_1 for energies $E = 35, 20,$ and 15 (at the lower boundary of the upper allowed region of Fig. 1). *Lower row*: trajectories *outside* the unit disk \mathcal{D}_1 , for energies $E = 3$ (at the upper boundary of the lower allowed region), $1,$ and -0.9 (near to the lower forbidden region), which fall completely outside the disk and correspond to the hyperbolic case to be seen in Sec. IV. We mark the scale 1 on the y -axis, understood to be in units of $1/\sqrt{|\alpha|}$.

The trajectories are circular when $\varepsilon = 0$, i.e., $C = 0$ or $E + \alpha p_\phi^2 = \pm |\beta p_\phi|$. This is the case of the upper right and lower left trajectories in Fig. 2. For $\alpha < 0$, it occurs on the two parabolas that bound the region excluded by the first condition in (22) and respect the other two inequalities. The radius of those circles can be found from (21), as $r^2(\phi) = D$. At the upper boundary, one has $E = -\alpha p_\phi^2 + |\beta p_\phi| \geq -\alpha p_\phi^2$, so in the Zernike $\alpha < 0$ region this means $E \geq |\alpha| p_\phi^2$, which in turn entails that $|\alpha| B \leq A$ or $D \leq 1/|\alpha|$, which yields the radius of the circle as $r_o = 1/\sqrt{|\alpha|}$; in the case $\alpha_Z = -1$ this is the boundary of the unit circle of Zernike’s differential equation.²⁷ On the other hand, at the lower boundary in the same Zernike range $\alpha < 0$, $E = |\alpha| p_\phi^2 - |\beta p_\phi|$, and one has $r_o'^2 = D = 2p_\phi^2 / (2|\alpha| p_\phi^2 - |\beta p_\phi|) > 1/|\alpha|$, which for $\alpha_Z = -1$ exceeds the unit radius allotted by Zernike’s requirement. We conclude that the elliptic trajectories in the lower “allowed” region of Fig. 1 (left) cannot correspond with solutions of the Zernike differential equation (1). Only those in the upper region do. On the other extreme of the $\alpha < 0$ region, the trajectories become lines when $\varepsilon \rightarrow 1$, namely, for every larger E and also when E approaches the lower boundary $-\beta^2/4|\alpha|$.

Regarding the region $\alpha > 0$ in Fig. 1 (right), the excentricity in (21) is $\varepsilon = 0$ on the parabola $E = -\alpha p_\phi^2 + |\beta p_\phi|$. The radii of those circles can be found as we did above, yielding $r_o^2(\phi) = 2p_\phi^2 / (|\beta p_\phi| - 2\alpha p_\phi^2)$. The trajectory is a *unit* circle when $2(1 + \alpha)p_\phi^2 = |\beta p_\phi|$, i.e., $|p_\phi| = |\beta|/2(\alpha + 1) < |\beta|/2\alpha$. This value falls on a single point of the parabolic boundary of the allowed region in Fig. 1 (right). On the upper boundary of that region, $E = \beta^2/4\alpha$, the excentricity is $\varepsilon = 1$ and the trajectories are lines. Finally, when $\alpha = 0$ and the allowed region is $E \geq |\beta p_\phi|$, on its boundary we have $\varepsilon = 0$ circles of radii $r_o^2 = 2|p_\phi|/\beta$.

IV. SEPARATION OF VARIABLES AND SYMMETRIES

The classical Zernike Hamiltonian (4) in Cartesian coordinates can be subject to the Hamilton-Jacobi method of solution with the action partial derivatives $p_x = \partial S/\partial x$ and $p_y = \partial S/\partial y$ and yields

the Hamiltonian (4) written as

$$H = \left(\frac{\partial S}{\partial x}\right)^2 + \left(\frac{\partial S}{\partial y}\right)^2 + \alpha \left(x \frac{\partial S}{\partial x} + y \frac{\partial S}{\partial y}\right)^2 - i\beta \left(x \frac{\partial S}{\partial x} + y \frac{\partial S}{\partial y}\right) = E. \quad (35)$$

This equation is separable on the (x, y) -plane, but the boundary condition imposed by Zernike²⁷ on the solutions, namely, that their absolute value at the boundary $x^2 + y^2 = 1$ be constant, can only be separated in polar coordinates, as we did in Sec. II. Although the classical Zernike system appears to belong to the class of Bertrand systems³ in which all bounded orbits are closed, it does not qualify as such because the linear and quadratic $\mathbf{r} \cdot \nabla$ terms replace the two-dimensional central force potentials of the Coulomb or isotropic oscillator systems. We surmise that this feature is a specific consequence of the superintegrability of the Zernike system. It is therefore of interest to find any additional separable systems of orthogonal coordinates and, associated with these, the extra symmetry operators that will clearly demonstrate the classical Zernike Hamiltonian to be superintegrable. We remind the reader that in an N -dimensional space with constant curvature (real or complex), a maximally superintegrable system allows, in addition to the Hamiltonian H , another $2N - 2$ functionally independent constants of motion, $L_1, L_2, \dots, L_{2N-2}, L_{2N-1} := H$, which are in involution with H , namely, $\{H, L_i\} = 0$ for $i \in \{1, 2, \dots, 2N - 2\}$.¹⁴

A. Coordinate systems on sphere and hyperboloid

Equation (1) is linear and of second order,

$$(1 + \alpha x^2) \frac{\partial^2 f}{\partial x^2} + 2\alpha xy \frac{\partial^2 f}{\partial x \partial y} + (1 + \alpha y^2) \frac{\partial^2 f}{\partial y^2} + (\alpha + \beta) \left(x \frac{\partial f}{\partial x} + y \frac{\partial f}{\partial y}\right) = -Ef. \quad (36)$$

According to the standard classification, this equation is of elliptic type when $-\alpha r^2 < 1$, parabolic type when $-\alpha r^2 = 1$, and hyperbolic type when $-\alpha r^2 > 1$. The original Zernike case $\alpha_Z = -1$ is in the range $\alpha < 0$, where the region of ellipticity is the interior of the circle $r < 1/\sqrt{|\alpha|}$. On the other hand, when $\alpha \geq 0$, the Eq. (1) is of elliptic type over the whole x - y plane \mathcal{R}^2 .

To be within the Zernike case, we consider first the range $\alpha < 0$ and map the open disk $x^2 + y^2 < 1/|\alpha| =: R^2$ on the hemisphere $\xi_1^2 + \xi_2^2 + \xi_3^2 = R^2$, $\xi_3 \geq 0$, embedded in a Euclidean space with three Cartesian coordinates ξ_i , through the orthogonal (or “vertical”) projection

$$\xi_1 = x, \quad \xi_2 = y, \quad \xi_3 = \sqrt{R^2 - x^2 - y^2}. \quad (37)$$

In these coordinates, the Hamiltonian equation (35) can be separated into three mutually orthogonal spherical systems of coordinates,¹⁷

System I:

$$\xi_1 = R \sin \vartheta \cos \varphi, \quad \xi_2 = R \sin \vartheta \sin \varphi, \quad \xi_3 = R \cos \vartheta, \quad \vartheta|_0^{\pi/2}, \varphi|_0^{2\pi}, \quad (38)$$

System II:

$$\xi_1 = R \cos \vartheta, \quad \xi_2 = R \sin \vartheta \cos \varphi, \quad \xi_3 = R \sin \vartheta \sin \varphi, \quad \vartheta|_0^\pi, \varphi|_0^\pi, \quad (39)$$

System III:

$$\xi_1 = R \sin \vartheta \sin \varphi, \quad \xi_2 = R \cos \vartheta, \quad \xi_3 = R \sin \vartheta \cos \varphi, \quad \vartheta|_0^\pi, \varphi|_{-\frac{1}{2}\pi}^{\frac{1}{2}\pi}, \quad (40)$$

and in the *elliptical* system of coordinates^{12,13,17} to be seen below.

Still within the $\alpha < 0$ case, we can consider the *outside* of the circle at radii $r^2 > 1/|\alpha|$, where the Eq. (36) is hyperbolic. There one can map the trajectories of the x - y plane on trajectories on the one-sheeted half-hyperboloid $\xi_1^2 + \xi_2^2 - \xi_3^2 = R^2 = 1/|\alpha|$. Coordinates that permit separation of variables for (36) replace trigonometric functions by hyperbolic functions thus

System H'I (pseudo-spherical):

$$\xi_1 = R \cosh \tau \cos \varphi, \quad \xi_2 = R \cosh \tau \sin \varphi, \quad \xi_3 = R \sinh \tau, \quad \tau \in \mathcal{R}, \varphi|_0^{2\pi}, \quad (41)$$

System H'II (equidistant):

$$\xi_1 = \pm R \cosh \tau_1, \quad \xi_2 = R \sinh \tau_1 \sinh \tau_2, \quad \xi_3 = R \sinh \tau_1 \cosh \tau_2, \quad \tau_1, \tau_2 \in \mathcal{R}, \quad (42)$$

System H'III (equidistant):

$$\xi_1 = R \cosh \tau \sin \varphi, \quad \xi_2 = R \cos \varphi, \quad \xi_3 = R \sinh \tau \sin \varphi, \quad \tau \in \mathcal{R}, \quad \varphi \in [0, 2\pi]. \quad (43)$$

On the other hand when $\alpha > 0$, the region of ellipticity being the whole plane \mathcal{R}^2 , allows one to map this plane on the upper sheet of the two-sheeted hyperboloid $\xi_3^2 - \xi_1^2 - \xi_2^2 = \varrho^2 = 1/\alpha$ using “modified” coordinate systems^{23,24}

System HI (pseudo-spherical):

$$\xi_1 = \varrho \sinh \tau \cos \varphi, \quad \xi_2 = \varrho \sinh \tau \sin \varphi, \quad \xi_3 = \varrho \cosh \tau, \quad \tau \in \mathcal{R}, \quad \varphi \in [0, 2\pi], \quad (44)$$

System HII (equidistant):

$$\xi_1 = \varrho \sinh \tau_1, \quad \xi_2 = \varrho \cosh \tau_1 \sinh \tau_2, \quad \xi_3 = \varrho \cosh \tau_1 \cosh \tau_2, \quad \tau_1, \tau_2 \in \mathcal{R}, \quad (45)$$

System HIII (equidistant):

$$\xi_1 = \varrho \cosh \tau'_1 \sinh \tau'_2, \quad \xi_2 = \varrho \sinh \tau'_1, \quad \xi_3 = \varrho \cosh \tau'_1 \cosh \tau'_2, \quad \tau'_1, \tau'_2 \in \mathcal{R}. \quad (46)$$

The hyperboloidal coordinates in (41)–(46) have been defined in Ref. 18.

B. Separation in spherical systems I, H'I, and HI

In the spherical coordinates (ϑ, φ) of system I in (38) for $\alpha < 0$, the Hamilton-Jacobi expression in (35) acquires the form

$$\frac{1 + \alpha R^2 \sin^2 \vartheta}{R^2 \cos^2 \vartheta} \left(\frac{\partial S}{\partial \vartheta} \right)^2 - i\beta \tan \vartheta \left(\frac{\partial S}{\partial \vartheta} \right) + \frac{1}{R^2 \sin^2 \vartheta} \left(\frac{\partial S}{\partial \varphi} \right)^2 = E. \quad (47)$$

This equation is integrable with the help of the first-order integral of motion

$$I_1 := p_\varphi = xp_y - yp_x, \quad (48)$$

that is independent of (α, β) and separates the action function as $S(\vartheta, \varphi) = S_1(\vartheta) + p_\varphi \varphi$, leading to the equation

$$\frac{1 + \alpha R^2 \sin^2 \vartheta}{R^2 \cos^2 \vartheta} \left(\frac{dS_1}{d\vartheta} \right)^2 - i\beta \tan \vartheta \left(\frac{dS_1}{d\vartheta} \right) + \frac{p_\varphi^2}{R^2 \sin^2 \vartheta} = E. \quad (49)$$

Using the same approach of Sec. III for the Zernike $\alpha < 0$ case, one finds the trajectory $\vartheta(\varphi)$ to be

$$\sin^2 \vartheta = \frac{|\alpha| D}{1 - \varepsilon \cos 2\varphi}, \quad (50)$$

where D and ε are given in (21), and which lies within the hemisphere of radius $R = 1/\sqrt{|\alpha|}$, as seen in Fig. 3. The trajectories reach the rim $\vartheta = \frac{1}{2}\pi$ only when $\beta p_\varphi = 0$.

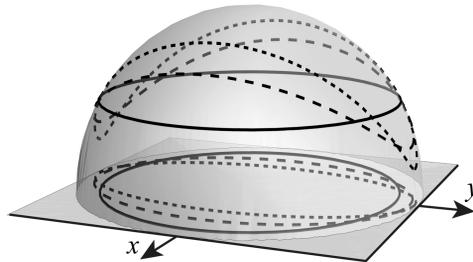


FIG. 3. Trajectories $\vartheta(\varphi)$ on the hemisphere given by (50) for the allowed upper regions of the Zernike system $(\alpha_Z, \beta_Z) = (-1, -2)$ in Fig. 1 (left) and their projection on the x - y plane inside the unit disk \mathcal{D}_1 , for the values corresponding to the upper row of orbits in Fig. 2: $p_\varphi = 3$ and energies $E = 15$ (continuous line, the circular orbit at the boundary of the allowed region), $E = 20$ (dashed line), and $E = 35$ (dotted line).

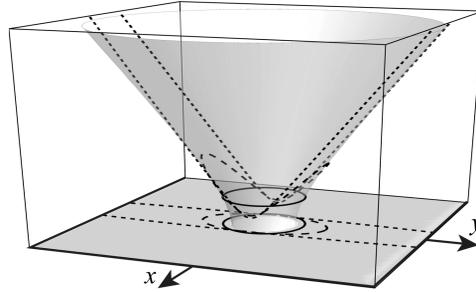


FIG. 4. Trajectories $\tau(\varphi)$ on the half-hyperboloid (of one sheet) given by (52), with $\alpha = -1$ and $\beta = -2$, and their projection on the x - y plane outside the unit disk \mathcal{D}_1 . The parameter values are the same as in the second row of Fig. 2, namely, $p_\phi = 3$ and energies $E = 3$ (at the upper boundary of the lower allowed region, marked by a continuous line), 1 (dashed line), and -0.9 (near to the lower forbidden region, dotted line).

Still in the $\alpha < 0$ case, the pseudo-spherical coordinates (τ, φ) of system H'I in (41) allow separation of the action function as $S(\tau, \varphi) = S_1(\tau) + p_\varphi \varphi$, so the Hamiltonian (35) leads to the equation

$$\alpha \left(\frac{dS_1}{d\tau} \right)^2 - i\beta \coth \tau \left(\frac{dS_1}{d\tau} \right) - \frac{\alpha p_\varphi^2}{\cosh^2 \tau} = E. \tag{51}$$

Then the trajectories, instead of (50), are given by

$$\cosh^2 \tau = \frac{|\alpha| D}{1 - \varepsilon \cos 2\varphi}, \tag{52}$$

with D and ε given in (21). These are closed orbits in the region $r^2 > 1/|\alpha|$. In Fig. 4, we show such trajectories on the one-sheeted half-hyperboloid.

Turning now to the case $\alpha > 0$ for the pseudo-spherical system (44), the separation of variables $S(\tau, \varphi) = S_1(\tau) + p_\varphi \varphi$ yields

$$\frac{1 + \alpha \varrho^2 \sinh^2 \tau}{\varrho^2 \cosh^2 \tau} \left(\frac{dS_1}{d\tau} \right)^2 - i\beta \tanh \tau \left(\frac{dS_1}{d\tau} \right) + \frac{p_\varphi^2}{\varrho^2 \sinh^2 \tau} = E, \tag{53}$$

so that the trajectory $\vartheta(\varphi)$ is found as

$$\sinh^2 \tau = \frac{\alpha D}{1 - \varepsilon \cos 2\varphi}, \tag{54}$$

lying on one sheet of a two-sheeted hyperboloid $\varrho^2 = 1/\alpha$ and where again D and ε are given in (21). The orbits on this manifold are elliptic and are shown in Fig. 5.

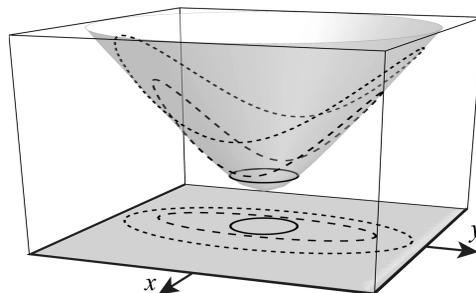


FIG. 5. Trajectories $\tau(\varphi)$ on the lower half-hyperboloid (of two sheets) given by (54) with $\alpha = +1$ and $\beta = -2$ and their projection on the full x - y plane. The parameter values are all near to the cusp of the allowed region in Fig. 1 (right): $p_\phi = 0.5$, $E = 0.75$ (continuous line); $p_\phi = 0.75$, $E = 0.97$ (dashed line); $p_\phi = 0.9$, $E = 0.993$ (dotted line).

C. Separation in coordinate systems II and III

The second system of spherical coordinates (ϑ, φ) in (39) leads to the Hamiltonian (35) in the form

$$\begin{aligned} & \frac{1 + \alpha R^2 \cos^2 \vartheta}{R^2 \sin^2 \vartheta} \left(\frac{\partial S}{\partial \vartheta} \right)^2 + i\beta \cot \vartheta \left(\frac{\partial S}{\partial \vartheta} \right) \\ & + \frac{1}{R^2 \sin^2 \vartheta} \left[1 + \frac{(1 + \alpha R^2) \cot^2 \varphi}{\sin^2 \vartheta} \left(\frac{\partial S}{\partial \varphi} \right)^2 \right] + i\beta \frac{\cot \varphi}{\sin^2 \vartheta} \left(\frac{\partial S}{\partial \varphi} \right) \\ & + 2 \frac{1 + \alpha R^2}{R^2 \sin^2 \vartheta} \cot \vartheta \left(\frac{\partial S}{\partial \vartheta} \right) \cot \varphi \left(\frac{\partial S}{\partial \varphi} \right) = E. \end{aligned} \quad (55)$$

When $\alpha < 0$, separation of variables applies on the action function $S(\vartheta, \varphi) = S_1(\vartheta) + S_2(\varphi)$ and leads to the pair of equations

$$-\alpha \left(\frac{dS_1}{d\vartheta} \right)^2 + i\beta \cot \vartheta \left(\frac{dS_1}{d\vartheta} \right) + \frac{K_{\text{II}}^2}{\sin^2 \vartheta} = E, \quad (56)$$

$$-\alpha \left(\frac{dS_2}{d\varphi} \right)^2 + i\beta \cot \varphi \left(\frac{dS_2}{d\varphi} \right) = K_{\text{II}}^2, \quad (57)$$

where K_{II}^2 is a separation constant. Rewriting (57) in Cartesian (x, y) coordinates, we obtain

$$\left[1 + \alpha(x^2 + y^2) \right] \left(\frac{dS_2}{dy} \right)^2 - i\beta y \left(\frac{dS_2}{dy} \right) = K_{\text{II}}^2. \quad (58)$$

The integration in y yields a second integral of motion that depends on the parameters (α, β) ,

$$I_2 := K_{\text{II}}^2 = \left[1 + \alpha(x^2 + y^2) \right] p_y^2 - i\beta y p_y. \quad (59)$$

In the case $\alpha > 0$, the action function admits separation of variables in the hyperbolic equidistant system III in (42), $S(\tau_1, \tau_2) = S_1(\tau_1) + S_2(\tau_2)$ and yields the two equations

$$\alpha \left(\frac{dS_2}{d\tau_2} \right)^2 - i\beta \tanh \tau_2 \left(\frac{dS_2}{d\tau_2} \right) = K_{\text{III}}^2, \quad (60)$$

$$\alpha \left(\frac{dS_1}{d\tau_1} \right)^2 - i\beta \tanh \tau_1 \left(\frac{dS_1}{d\tau_1} \right) + \frac{K_{\text{III}}^2}{\cosh^2 \tau_1} = E, \quad (61)$$

which lead to the same integral of motion I_2 in (59).

D. Separation in the coordinate system III

The third spherical system of coordinates in (40) leads to the Hamilton-Jacobi form (35) written as

$$\begin{aligned} & \frac{1 + \alpha R^2 \cos^2 \vartheta}{R^2 \sin^2 \vartheta} \left(\frac{\partial S}{\partial \vartheta} \right)^2 + i\beta \cot \vartheta \left(\frac{\partial S}{\partial \vartheta} \right) \\ & + \frac{1}{R^2 \sin^2 \vartheta} \left[1 + \frac{(1 + \alpha R^2) \tan^2 \varphi}{\sin^2 \vartheta} \left(\frac{\partial S}{\partial \varphi} \right)^2 \right] - i\beta \frac{\tan \varphi}{\sin^2 \vartheta} \left(\frac{\partial S}{\partial \varphi} \right) \\ & - 2 \frac{1 + \alpha R^2}{R^2 \sin^2 \vartheta} \cot \vartheta \left(\frac{\partial S}{\partial \vartheta} \right) \tan \varphi \left(\frac{\partial S}{\partial \varphi} \right) = E. \end{aligned} \quad (62)$$

In the case $\alpha < 0$, for $R^2 = -1/\alpha$, the separation of variables in the action function, $S(\vartheta, \varphi) = S_3(\vartheta) + S_4(\varphi)$ leads to

$$-\alpha \left(\frac{dS_3}{d\vartheta} \right)^2 + i\beta \cot \vartheta \left(\frac{dS_3}{d\vartheta} \right) + \frac{K_{\text{III}}^2}{\sin^2 \vartheta} = E, \quad (63)$$

$$-\alpha \left(\frac{dS_4}{d\varphi} \right)^2 - i\beta \tan \varphi \left(\frac{dS_4}{d\varphi} \right) = K_{\text{III}}^2, \quad (64)$$

From (64), we find a third constant of motion that depends on (α, β) ,

$$I_3 := K_{III}^2 = \left[1 + \alpha(x^2 + y^2) \right] p_x^2 - i\beta x p_x, \tag{65}$$

and which under the phase space $\frac{1}{2}\pi$ -rotation $(x, p_x; y, p_y) \leftrightarrow (y, p_y; -x, -p_x)$ coincides with I_2 in (59). Finally, we note that when $\alpha > 0$, the separations of variables (44)–(46) on the hyperboloid yield the same integrals of motion I_1, I_2 , and I_3 given above.

We note that unlike the three orthogonal coordinate systems on the sphere, on hyperboloids there are *nine* orthogonal coordinate systems where the Laplace and the Helmholtz equations yield to separation of variables.¹⁶

E. Separation of variables in the elliptic system

The Hamilton-Jacobi equation (35) also yields to separation in *elliptic* coordinates on the sphere in trigonometric form,^{12,13,17}

$$\xi_1^e = R \cos \varphi \sqrt{1 - k_1^2 \cos^2 \vartheta}, \quad \xi_2^e = R \sin \vartheta \sin \varphi, \quad \xi_3^e = R \cos \vartheta \sqrt{1 - k_3^2 \cos^2 \varphi}, \tag{66}$$

where the constants $k_1 := \cos f$ and $k_3 := \sin f$ are related to the interfocal distance $2f$ of the ellipses on the upper unit hemisphere, so that $k_1^2 + k_3^2 = 1$. When $\alpha < 0$ and thus $R^2 = 1/|\alpha|$, the action function separates as $S(\vartheta, \varphi) = S_1^e(\vartheta) + S_2^e(\varphi)$ and leads again to two equations,

$$(1 - k_1^2 \cos^2 \vartheta) \left[\alpha S_1^{e\prime 2} + i\beta \tan \vartheta S_1^{e\prime} \right] + Ek_1^2 \sin^2 \vartheta = K_e^2, \tag{67}$$

$$\alpha(1 - k_3^2 \cos^2 \varphi) S_2^{e\prime 2} - i\beta k_3^2 \cos \varphi \sin \varphi S_2^{e\prime} + Ek_3^2 \sin^2 \varphi = -K_e^2, \tag{68}$$

where K_e^2 is a separation constant, $S_1^{e\prime} := dS_1/d\vartheta$ and $S_2^{e\prime} := dS_2/d\varphi$. Eliminating E from these equations, one obtains

$$\begin{aligned} & K_e^2(k_3^2 \sin^2 \varphi + k_1^2 \sin^2 \vartheta) \\ &= k_3^2 \sin^2 \varphi (1 - k_1^2 \cos^2 \vartheta) \left[\alpha (S_1^{e\prime})^2 + i\beta \tan \vartheta S_1^{e\prime} \right] \\ & \quad - k_1^2 \sin^2 \vartheta \left[\alpha (1 - k_3^2 \cos^2 \varphi) (S_2^{e\prime})^2 - i\beta k_3^2 \cos \varphi \sin \varphi S_2^{e\prime} \right]. \end{aligned} \tag{69}$$

Returning to Cartesian (x, y) coordinates,

$$\frac{S_1^{e\prime}}{R} = p_x \frac{k_1^2 \sin \vartheta \cos \vartheta \cos \varphi}{\sqrt{1 - k_1^2 \cos^2 \vartheta}} + p_y \cos \vartheta \sin \varphi, \tag{70}$$

$$\frac{S_2^{e\prime}}{R} = -p_x \sqrt{1 - k_1^2 \cos^2 \vartheta} \sin \varphi + p_y \sin \vartheta \cos \varphi, \tag{71}$$

we can express the constant K_e^2 as

$$\begin{aligned} K_e^2 &= -\alpha k_1^2 y^2 p_x^2 + 2\alpha k_1^2 x y p_x p_y \\ & \quad - \left(\alpha k_1^2 x^2 + k_3^2 [1 + \alpha(x^2 + y^2)] \right) p_y^2 + i\beta k_3^2 y p_y \\ &= -\alpha k_1^2 I_1^2 - k_3^2 I_2. \end{aligned} \tag{72}$$

Thus the elliptic separation constant K_e^2 is not functionally independent but depends on the constants I_1 and I_2 in (48) and (59).

V. ALGEBRAIC STRUCTURE AND CONCLUSIONS

We have found three functionally independent integrals of motion, I_1 in (48), I_2 in (59), and I_3 in (65), with no singularities on the full (α, β) parameter space. To probe their algebraic structure, let us define

$$J_1 := \frac{1}{2}I_1 = \frac{1}{2}(xp_y - yp_x), \quad (73)$$

$$J_2 := \frac{1}{2}(I_3 - I_2) = \frac{1}{2} \left[1 + \alpha(x^2 + y^2) \right] (p_x^2 - p_y^2) - \frac{1}{2}i\beta(xp_x - yp_y). \quad (74)$$

The function J_1 is $\frac{1}{2}$ -angular momentum and its Poisson operator $\{J_1, \circ\}$ generates rotations of phase space, while the function J_2 does depend on (α, β) . These functions Poisson-commute with the Zernike Hamiltonian function $H^{(\alpha, \beta)}$ in (4), which can be written as

$$H^{(\alpha, \beta)} = I_3 + I_2 - \alpha I_1^2, \quad (75)$$

but do not commute with each other. This shows that the generalized classical (α, β) -Hamiltonian of Zernike, $H^{(\alpha, \beta)}$ in (5), is superintegrable on each of the domains examined above, in particular on the (x, y) -disk \mathcal{D}_R , $r < R = 1/\sqrt{|\alpha|}$ for $\alpha < 0$, that contains the Zernike original case $(\alpha_Z, \beta_Z) = (-1, -2)$.

To identify the symmetry of the generalized Zernike (α, β) -Hamiltonians, we introduce a new integral of motion through the Poisson bracket of (73) and (74),

$$J_3 := \{J_1, J_2\} = \left[1 + \alpha(x^2 + y^2) \right] p_x p_y - i\frac{1}{2}\beta(xp_y + yp_x), \quad (76)$$

which also Poisson-commutes with $H^{(\alpha, \beta)}$ and is functionally independent of J_1 and J_2 , although it can be seen that J_2 and J_3 are connected to each other by a rotation of $\frac{1}{4}\pi$ in the x - y phase space planes. The algebraic structure of three functions J_1, J_2, J_3 is thus found to be

$$\{J_3, J_1\} = J_2, \quad \{J_1, J_2\} = J_3, \quad (77)$$

$$\{J_2, J_3\} = J_1(\beta^2 - 2\alpha H^{(\alpha, \beta)} - 8\alpha^2 J_1^2). \quad (78)$$

They form therefore a cubic Higgs algebra⁹ that Poisson-commutes with the generalized Zernike Hamiltonian, $\{J_i, H^{(\alpha, \beta)}\} = 0$.

When $\alpha \rightarrow 0$ so $R \rightarrow \infty$, the Zernike Hamiltonian becomes a simpler quadratic function,

$$H^{(0, \beta)}(\mathbf{q}, \mathbf{p}) = \mathbf{p}^2 - i\beta \mathbf{q} \cdot \mathbf{p}. \quad (79)$$

The Poisson operators of all quadratic functions of these four phase space coordinates close under commutation into the real symplectic Lie algebra $\mathfrak{sp}(4, \mathbb{R})$.

The Hamiltonian (79) belongs to the elliptic orbit of harmonic oscillators (Ref. 25, Chap. 12), as can be seen under the complex linear canonical transformation

$$\begin{pmatrix} \mathbf{p} \\ \mathbf{q} \end{pmatrix} = \begin{pmatrix} \mathbf{1}/\sqrt{2} & i\beta\mathbf{1}/\sqrt{2} \\ \mathbf{0} & \sqrt{2}\mathbf{1} \end{pmatrix} \begin{pmatrix} \mathbf{P} \\ \mathbf{Q} \end{pmatrix}. \quad (80)$$

This maps (79) on a regular harmonic oscillator,

$$F_0 := H^{(0, \beta)}(\mathbf{Q}, \mathbf{P}) = \frac{1}{2}(\mathbf{P}^2 + \beta^2 \mathbf{Q}^2), \quad (81)$$

and the three constants of the motion, J_1, J_2, J_3 in (73), (74), and (76), on

$$F_1 := \frac{1}{2} \mathbf{Q} \times \mathbf{P} = \frac{1}{2}(Q_x P_y - Q_y P_x), \quad (82)$$

$$F_2 := \frac{1}{2}(P_x^2 + \beta Q_x^2) - \frac{1}{2}(P_y^2 + \beta^2 Q_y^2), \quad (83)$$

$$F_3 := \frac{1}{2}P_x P_y + \frac{1}{2}\beta^2 Q_x Q_y, \quad (84)$$

whose Poisson brackets close into a scaled $\mathfrak{u}(2)$ Lie algebra,

$$\{F_1, F_2\} = F_3, \quad \{F_2, F_3\} = \beta^2 F_1, \quad \{F_3, F_1\} = F_2, \quad \{F_0, F_i\} = 0. \quad (85)$$

In the paraxial geometric or wave optical interpretation, the central $F_0 \in \mathfrak{u}(1)$ generates isotropic fractional Fourier transforms,²¹ while F_2 generates anisotropic ones, F_1 generates rotations, and F_3 generates gyrations¹⁹ that transform Hermite-Gauss into Laguerre-Gauss beams. Together their

Poisson operators form the *Fourier algebra*,²¹ which is the maximal compact subalgebra in $\mathfrak{sp}(4, \mathbb{R})$. If β were a pure imaginary number, (81) would be the repulsive oscillator Hamiltonian and (82)–(84) its commuting “Fourier” algebra $\mathfrak{su}(1, 1) = \mathfrak{so}(2, 1)$; a similar treatment of the classical system with Hamiltonian (4) would yield hyperbolic orbits. For $\beta = 0$, a free system with an inhomogeneous $\mathfrak{iso}(2)$ “Fourier” algebra would appear.

The original Zernike system $\hat{Z}^{(\alpha_Z, \beta_Z)}$ in (1)²⁷ was proposed to develop a set orthogonal and a complete set of two-variable orthogonal polynomials $Z_{n,m}(r) \exp(im\phi)$, $Z_{n,m}(1) = 1$, $|m| \leq n$, which present the same (n, m) -pattern as the two-dimensional quantum harmonic oscillator states. There has been some effort in replicating the raising and lowering techniques of the oscillator scheme on the Zernike system^{20,26} without achieving a proper Lie algebra. Because here we have a two-parameter system $H^{(\alpha, \beta)}$, we could surmise that superintegrable systems can be obtained as a new kind of algebra deformation, from (81)–(85) to (73)–(78), consisting in the addition of the square of an element of a Lie algebra to the generator designed to be the original quadratic Hamiltonian. Imposing boundary conditions such as those proposed by Zernike will need the quantum treatment of this construction.

ACKNOWLEDGMENTS

We acknowledge the interest and early discussions with Professor Natig M. Atakishiyev (Instituto de Matemáticas, UNAM); we thank Guillermo Krötzsch (ICFUNAM) for indispensable help with the figures. G.S.P. and A.Y. thank the support of project PROSNI-2016 (Universidad de Guadalajara). K.B.W. thanks Cristina Salto-Alegre (Posgrado en Ciencias Físicas, ICFUNAM) for her interest and interaction on the subject, and acknowledges the support of UNAMDGAPA Project *Óptica Matemática* PAPIIT-IN101115.

- ¹ *Non-Selfadjoint Operators in Quantum Physics: Mathematical Aspects*, edited by Bagarello, F., Gazeau, J.-P., Szafraniec, F. H., and Znojil, M. (John Wiley & Sons, Hoboken, 2015).
- ² Bender, C. M., Brody, D. C., and Müller, M. P., “Hamiltonian for the zeros of the Riemann zeta function,” *Phys. Rev. Lett.* **118**, 130201 (2017).
- ³ Bertrand, J., “Théorème relatif au mouvement d’un point attiré vers un centre fixe,” *C. R. Acad. Sci.* **77**, 849–853 (1873).
- ⁴ Bhatia, A. B. and Wolf, E., “On the circle polynomials of Zernike and related orthogonal sets,” *Math. Proc. Cambridge Philos. Soc.* **50**, 40–48 (1954).
- ⁵ Born, M. and Wolf, E., *Principles of Optics: Electromagnetic Theory of Propagation, Interference and Diffraction of Light*, 7th ed. (Cambridge University Press, 1999), p. 986.
- ⁶ Cariñena, J. F., Rañada, M. F., and Santander, M., “Two important examples of nonlinear oscillators,” e-print [arXiv:math-ph/0505028](https://arxiv.org/abs/math-ph/0505028).
- ⁷ Cariñena, J. F., Perelomov, A. M., and Rañada, M. F., “Isochronous classical systems and quantum systems with equally spaced spectra,” *J. Phys.: Conf. Ser.* **87**, 012007 (2007).
- ⁸ Gradshteyn, I. S. and Ryzhik, I. M., *Table of Integrals, Series, and Products*, 6th ed. (Academic Press, 2000).
- ⁹ Higgs, P. W., “Dynamical symmetries in a spherical geometry,” *J. Phys. A: Math. Gen.* **12**, 309–323 (1979).
- ¹⁰ Ismail, M. E. H. and Zhang, R., “Classes of bivariate orthogonal polynomials,” *SIGMA* **12**, 021 (2016); e-print [arXiv:1502.07256](https://arxiv.org/abs/1502.07256).
- ¹¹ Kintner, E. C., “On the mathematical properties of the Zernike polynomials,” *Opt. Acta* **23**, 679–680 (1976).
- ¹² Lukac, I. and Smorodinskiĭ, Ya. A., “Wave functions for the asymmetric top,” *Sov. Phys. JETP* **30**, 728–730 (1970).
- ¹³ Lukach, I., “A complete set of the quantum-mechanical observables on a two-dimensional sphere,” *Theor. Math. Phys.* **14**, 271–281 (1973).
- ¹⁴ Miller, Jr., W., Post, S., and Winternitz, P., “Classical and quantum superintegrability with applications,” *J. Phys. A: Math. Theor.* **47**, 423001 (2014).
- ¹⁵ Myrick, D. R., “A generalization of the radial polynomials of F. Zernike,” *SIAM J. Appl. Math.* **14**, 476–489 (1966).
- ¹⁶ Olevskiĭ, M. N., “Triorthogonal systems in spaces of constant curvature in which the equation $\Delta_2 u + \lambda u = 0$ allows a complete separation of variables,” *Mat. Sbornik* **27**, 379–426 (1950).
- ¹⁷ Pogosyan, G. S., Sissakian, A. N., and Winternitz, P., “Separation of variables and Lie algebra contractions. Applications to special functions,” *Phys. Part. Nuclei* **33**(Suppl. 1), S123–S144 (2002).
- ¹⁸ Pogosyan, G. S. and Yakhno, A., “Lie algebra contractions and separation of variables on two-dimensional hyperboloidal coordinate systems,” e-print [arXiv 1510.03785 V1](https://arxiv.org/abs/1510.03785) (2015).
- ¹⁹ Rodrigo, J. A., Alieva, T., and Bastiaans, T. J., “Phase space rotators and their applications in optics,” in *Optical and Digital Image Processing: Fundamentals and Applications*, edited by Cristóbal, G., Schelkens, P., and Thienpont, H. (Wiley-VCH Verlag, 2011), pp. 251–271.
- ²⁰ Shakibaei, B. H. and Paramesran, R., “Recursive formula to compute Zernike radial polynomials,” *Opt. Lett.* **38**, 2487–2489 (2013).
- ²¹ Simon, R. and Wolf, K. B., “Structure of the set of paraxial optical systems,” *J. Opt. Soc. Am. A* **17**, 342–355 (2000).

- ²² Tango, W. J., "The circle polynomials of Zernike and their application in optics," [Appl. Phys.](#) **13**, 327–332 (1977).
- ²³ Winternitz, P., Lukac, P., and Smorodinskiĭ, Ya. A., "Quantum numbers in the little groups of the Poincaré group," [Sov. J. Nucl. Phys.](#) **7**, 139–145 (1968).
- ²⁴ Winternitz, P., Smorodinsky, J., and Sheftel, M., "Poincaré and Lorentz invariant expansions of relativistic amplitudes," [Sov. J. Nucl. Phys.](#) **7**, 785–792 (1968).
- ²⁵ Wolf, K. B., *Geometric Optics on Phase Space* (Springer, Heidelberg, 2004).
- ²⁶ Wünsche, A., "Generalized Zernike or disc polynomials," [J. Comput. Appl. Math.](#) **174**, 135–163 (2005).
- ²⁷ Zernike, F., "Beugungstheorie des schneidenverfahrens und seiner verbesserten form der phasenkontrastmethode," [Physica](#) **1**, 689–704 (1934).